1 Problem set 2: Hints & Tipps

1.1 Formal solution of the radiative transfer equation

- re-use the class (or code) you wrote for the previous problem, in particular the tau grid and the infrastructure
- start by generating the points μ_j and weights w_j for the Gauss quadrature that replaces the J integration J(τ) = 1/2 ∑ w_jI(μ_j, τ). There's a numpy method for this:
 self.mu, self.w = np.polynomial.legendre.leggauss(nmu)
- You will need to write a routine that computes I(j, i) for all μ_j and τ_i
- This is done by compute I(j, :) for each j in succession by applying the scheme discussed in the lectures and the problem sheet.
- For this you need the coefficients α_i, β_i, γ_i. The method computeFScoeffs in the template does that for you, you need top provide S_{i-1}, S_i, S_{i+1} in Sm, S0, Sp and τ_{prev→curr} and τ_{curr→next} in dtau, dtau1. Set self.taulin=1e-2 and self.use_Bezier = False.
- It is easier to first (for a given direction μ_j) compute the α_i, β_i, γ_i for all i and store them in numpy arrays and then in a second step to compute the I(j, i) for all i and given j.
- Keep in mind that the direction of solving changes depending on the sign of μ_j . For $\mu > 0$ you need to got from large τ towards zero, for $\mu < 0$ it is from zero to τ_{max} . Similar for γ . Remember that the solution follows the propagation of the photons. α always considers the previous point, that can be either *i* or *i* + 1, depending on sign(μ_j).
- With the routine computing I(i, j), the formal solution is now easy, just call it for the different j and compute $J(i) = \sum_j w_j I(j, i)$. Note: you do NOT need to store I(j, i), just call your routine and the contributions for J on the fly. For this exercise it's easy to just store the full I(j, i), but imagine you have millions of τ points (3D models!) and thousands of j ...
- In general, it is best to scetch the solution on a (big) sheet of paper. Think first, code later. In a variation of a quote from Stanislav Lem: Think more, code less.